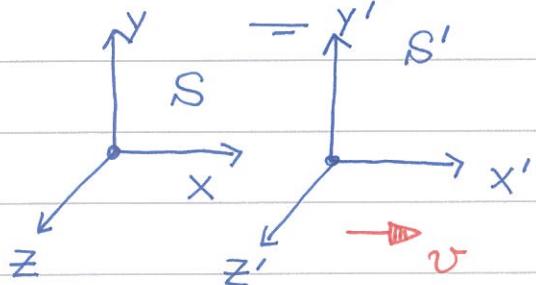




HH0082 Lorentz transformation

Newtonian mechanics is invariant under Galilean transformation,

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$$\begin{aligned}x' &= x - vt \\y' &= y \\z' &= z \\t' &= t\end{aligned}$$

absolute time.

Take the time derivative on both sides,

$$\frac{dx'}{dt} = \frac{dx}{dt} - v \rightarrow u'_x = u_x - v \text{ and also}$$

$$\begin{aligned}u'_y &= u_y \\u'_z &= u_z\end{aligned}$$

This is the velocity addition rule — our common intuitions.

Taking one more derivative gives

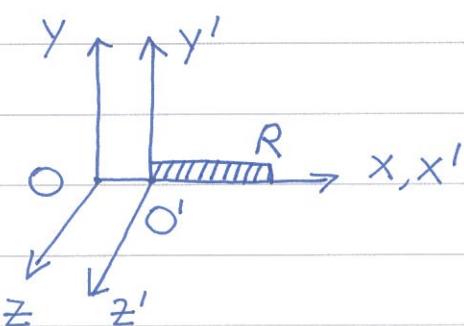
$$\frac{du'_x}{dt} = \frac{du_x}{dt}, \text{ similar identities for } y, z \text{ directions.}$$

In vector form, the result is very simple $\vec{a}' = \vec{a}$, i.e. the acceleration is invariant under Galilean transformation.

∅ Derivation of Lorentz transformation.

Let us ask Jack and Jill to help us again. A ruler

of natural length l_0 is moving with Jill. The origins of the inertia frames are chosen to be



$$(x, t) = (0, 0) = (x', t')$$

Jill's view

The ruler is at rest and the trajectory of the end point R is

$$x' = l_0 \text{ for all times } t'$$





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Jack's view

The length of the ruler is contracted to $\ell = \ell_0 \sqrt{1 - v^2/c^2}$ and the point R is moving at velocity v . Its trajectory is

$$x = \ell + vt = \ell_0 \sqrt{1 - v^2/c^2} + vt$$

Thus, we find the relation

$$x = x' \sqrt{1 - v^2/c^2} + vt \quad \text{OR}$$

$$x' = \gamma(x - vt)$$

Note that the principle of relativity requires that the form of the transformation from S to S' be identical to that from S' to S (except flipping v to $-v$). Thus, we expect

$$x = \gamma(x' + vt')$$

eliminate x'

$$t' = \gamma(t - \frac{v}{c^2}x)$$

Furthermore, because transverse length is invariant, $y' = y$ and $z' = z$. Collecting all relations together, we find the Lorentz transformation between (x, y, z, t) and (x', y', z', t') .

$$x' = \gamma(x - vt)$$

$$y' = y$$

$$z' = z$$

$$t' = \gamma(t - \frac{v}{c^2}x)$$

OR

$$x = \gamma(x' + vt')$$

$$y = y'$$

$$z = z'$$

$$t = \gamma(t' + \frac{v}{c^2}x')$$

The Lorentz transformation helps us to map events from S to S', or vice versa. For instance, we can understand length contraction and time dilation without "smart" arguments.

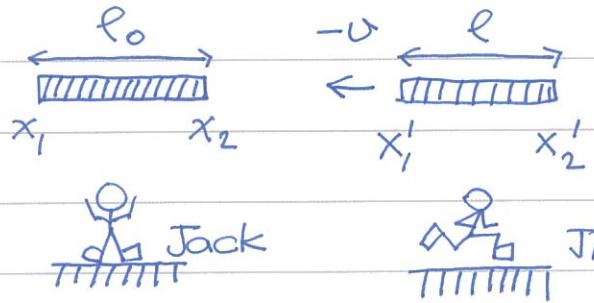
It is important to mention that Maxwell equations are invariant under Lorentz transformation.





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Length contraction revisited ☺



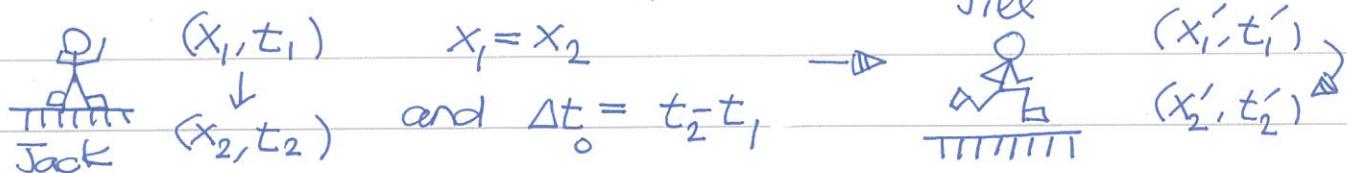
measured at the same time in Jill's frame, $t'_1 = t'_2$

$$l_0 = x_2 - x_1 = \gamma(x'_2 + vt'_2) - \gamma(x'_1 + vt'_1) = \gamma(x'_2 - x'_1) = \gamma l$$

Thus, $l = l_0 \sqrt{1 - v^2/c^2}$ as derived before by light clock.

time dilation revisited ☺

two events @ the same place.



The time interval in Jill's view is

$$t'_2 - t'_1 = \gamma(t_2 - \frac{v}{c^2}x_2) - \gamma(t_1 - \frac{v}{c^2}x_1) = \gamma(t_2 - t_1) = \gamma \Delta t_0$$

Thus, the time interval is dilated in the moving frame.

① Velocity addition rule: Let us make use of Lorentz transformation to see how velocities are related.

$$\begin{aligned} dx' &= \gamma(dx - vt dt) \\ dt' &= \gamma(dt - \frac{v}{c^2}dx) \end{aligned} \rightarrow u'_x = \frac{dx'}{dt'} = \frac{dx - vt dt}{dt - \frac{v}{c^2}dx}$$

$$\text{Thus, } u'_x = \frac{dx/dt - v}{1 - \frac{v}{c^2}(dx/dt)} \rightarrow u'_x = \frac{u_x - v}{1 - vu_x/c^2}$$

Special relativity gives rise to a non-trivial factor $(1 - vu_x/c^2)^{-1}$. You will see how this factor makes the speed of light constant later.





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velocity

addition

rule

Similarly, $dy' = dy$, $dz' = dz$.

$$u'_y = \frac{dy'}{dt'} = \frac{dy}{\gamma(dt - \frac{v}{c^2}dx)} = \frac{u_y}{\gamma(1 - v u_x / c^2)}$$

The same transformation for u'_z .

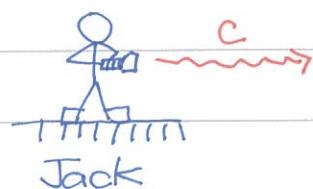
$$u'_x = \frac{u_x - v}{1 - v u_x / c^2}$$

$$u'_y = \frac{u_y}{\gamma(1 - v u_x / c^2)}$$

$$u'_z = \frac{u_z}{\gamma(1 - v u_x / c^2)}$$

Note that $u'_y \leftrightarrow u_y$ and $u'_z \leftrightarrow u_z$ are affected by the velocity component u_x — motion in x, y, z directions are not independent.

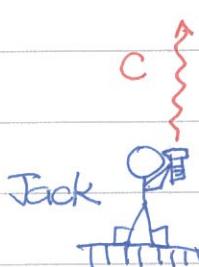
speed of light revisited. Consider a light propagating along the x axis. $\vec{u} = (u_x, u_y, u_z) = (c, 0, 0)$



$$u'_x = \frac{c - v}{1 - v/c} = c, \quad u'_y = 0 = u'_z$$

Thus, Jill will observe the same speed c .

What about propagating along the y axis?



$$\vec{u} = (0, c, 0)$$

$$u'_x = -v, \quad u'_y = \frac{c}{\gamma}, \quad u'_z = 0$$

Thus, the velocity is $\vec{u}' = (-v, \frac{c}{\gamma}, 0)$



The corresponding speed of light is

$$c' = \sqrt{u'_x^2 + u'_y^2 + u'_z^2} = \sqrt{v^2 + c^2(1 - v^2/c^2)} = c$$

One can see that the speed of light is invariant under Lorentz transformation. Very nice !!





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① Relativistic momentum.

In special relativity, we can show that the quantity

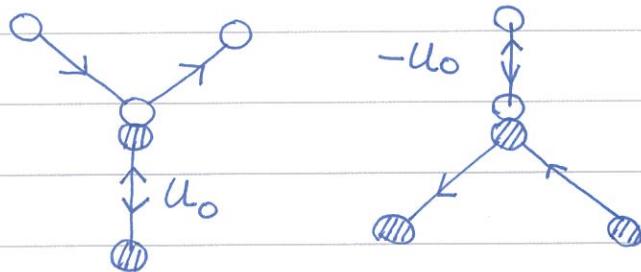
$$\vec{P} = \frac{1}{\sqrt{1-u^2/c^2}} m \vec{u}$$

is conserved in the elastic collision and thus should be the proper definition of momentum.

Consider the twins Jack and Jill throwing balls at each other to make a symmetric elastic collision as below.

Jack.Jill

◎ : Jack the same mass.
○ : Jill



Both of them throw out the balls vertically at velocity $(0, u_0, 0)$ Jack.
 $(0, -u_0, 0)$ Jill.

In Jill's view, the velocity of Jack's ball is

$$(0, u_0, 0) \rightarrow (-v, \sqrt{1-v^2/c^2} u_0, 0)$$

Assume the momentum acquires a correction factor that depends on the speed, $\vec{P} = \Gamma(u) m \vec{u}$

BEFORE collision.

$$P_y = -\Gamma(u_0) \cdot m u_0 + \Gamma(u_{\text{Jack}}) m \sqrt{1-u^2/c^2} u_0$$

$$\text{where the speed } u_{\text{Jack}} = \sqrt{v^2 + u_0^2(1-u^2/c^2)}.$$

AFTER collision

$$P'_y = \Gamma(u_0) m u_0 - \Gamma(u_{\text{Jack}}) m \sqrt{1-u^2/c^2} u_0$$

Momentum conservation requires $P_y = P'_y$ for the elastic collision considered here.





$$\rightarrow \Gamma(u_0) m u_0 = \Gamma(u_{\text{Jack}}) \sqrt{1 - u^2/c^2} m u_0$$

Now take $u_0 \rightarrow 0$ limit. The speed $u_{\text{Jack}} \rightarrow u$

$$\Gamma(0) = \Gamma(u) \sqrt{1 - u^2/c^2}, \text{ but we expect } \Gamma(0) = 1$$

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Thus, the prefactor is

The relativistic momentum

takes the form:

$$\Gamma(u) = \frac{1}{\sqrt{1 - u^2/c^2}} = \gamma$$

The correction becomes significant when the speed u is fast.

Newton's 2nd law in its most general form is

$$\vec{F} = \frac{d\vec{P}}{dt} = \frac{d}{dt} \left(\frac{m\vec{u}}{\sqrt{1 - u^2/c^2}} \right) \quad \text{valid relativistically.}$$

Some people like to introduce a relativistic mass

$$m_{\text{rel}} = \frac{m}{\sqrt{1 - u^2/c^2}} \rightarrow \infty, \text{ as } u \text{ approaches } c.$$

But it can cause confusions from time to time so I won't use the notion in my lectures.



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