

HHO133 -
SCHRODINGER
EQUATION



Schrodinger equation



The dynamics of the wave function is described by the **Schrodinger equation**. Note that the time derivative is only **first-order**.

$$-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \psi(x, t) + V(x) \psi(x, t) = i\hbar \frac{\partial}{\partial t} \psi(x, t)$$

Quantum Mechanics

To describe a quantum particle, one needs to know its Wave Function $\Psi(x, y, z, t)$.

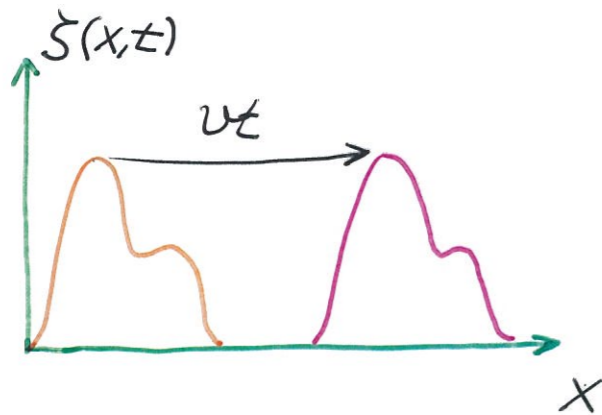
$$|\Psi(x, y, z, t)|^2 dx dy dz = \text{probability in } \begin{array}{c} dz \\ \text{cube} \\ dx \quad dy \end{array} \text{ at } (x, y, z)$$

To solve for the wave function $\Psi(x, y, z, t)$, one needs to understand Schrödinger Equation.

$$i\hbar \frac{\partial \Psi}{\partial t} = -\frac{\hbar^2}{2m} \left[\frac{\partial^2 \Psi}{\partial x^2} + \frac{\partial^2 \Psi}{\partial y^2} + \frac{\partial^2 \Psi}{\partial z^2} \right] + V \Psi$$

For simplicity, we would mainly concentrate on the 1D case.
Note that it is NOT the same as wave equation....

Compare with Wave Equation



A conventional wave obeys the following equation.

$$\frac{\partial^2 \xi}{\partial t^2} = v^2 \frac{\partial^2 \xi}{\partial x^2}$$

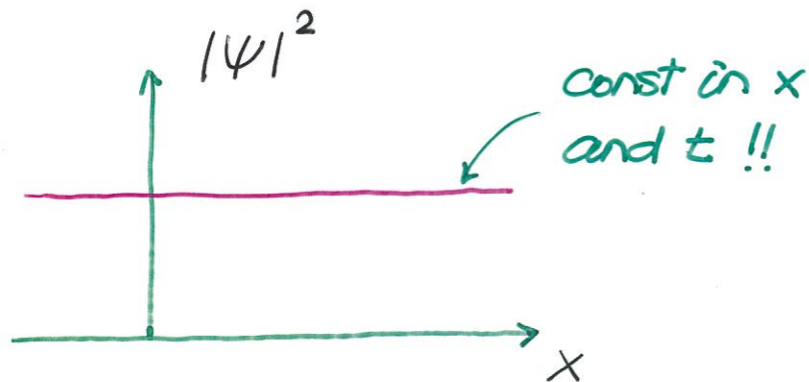
v : wave velocity.

On the other hand, the free particle satisfies the Schrödinger equation

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2}$$



$$\psi(x,t) = \frac{1}{\sqrt{L}} e^{i(kx - \omega t)}$$



The wave seems static if only $p(x,t)$ is measured.....

Time-Independent Schrödinger equation

Since the Schrödinger equation is first-order in time, its time dependence can be solved easily ☹️

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V \psi$$

educated guess !!

$$\psi(x,t) = \Phi(x) e^{-i\omega t}$$

$$i\hbar (-i\omega) \Phi e^{-i\omega t} = \left[-\frac{\hbar^2}{2m} \frac{\partial^2 \Phi}{\partial x^2} + V \Phi \right] e^{-i\omega t}, \quad \text{note that } E = \hbar\omega$$



$$-\frac{\hbar^2}{2m} \frac{d^2 \Phi}{dx^2} + V \Phi = E \Phi$$

☹️ should always keep in mind the t dependence !!

Or, sometimes in the form of

$$\frac{d^2 \Phi}{dx^2} + \frac{2m}{\hbar^2} (E - V) \Phi = 0$$

Q: Is E arbitrary?
Or.... NOT....

Revisit Free Particle.

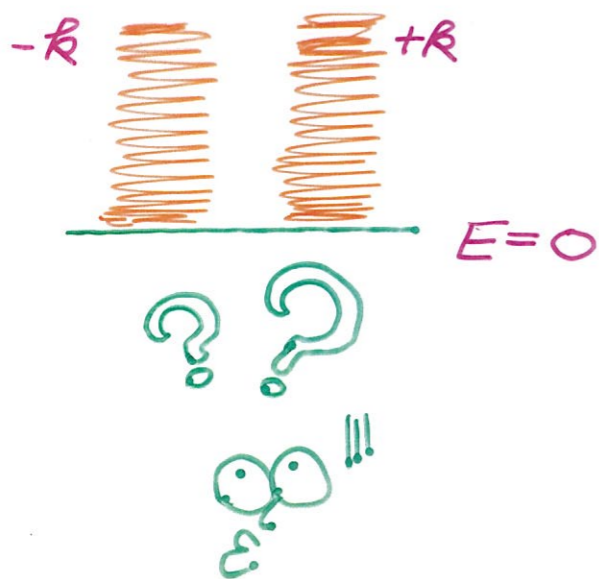
Since $V=0$, the equation is rather trivial.

$$\frac{d^2\psi}{dx^2} + \frac{2mE}{\hbar^2} \psi = 0 \quad \psi(x) = e^{iKx} \quad \text{with} \quad K^2 = \frac{2mE}{\hbar^2}$$

Or, in more familiar format $\frac{\hbar^2 K^2}{2m} = E \Rightarrow K = \pm \sqrt{\frac{2mE}{\hbar^2}}$

For convenience, $K = \sqrt{\frac{2mE}{\hbar^2}}$ and the two degenerate solutions are

$$\psi(x,t) = e^{-i\omega t} e^{\pm iKx}$$



As long as $E > 0$, there are pairs of solutions with the same energy.

- Q: (1) Why the 2-fold degeneracy?
(2) What happens if $E < 0$?

Time-Reversal Symmetry

Suppose a wave function $\psi(x,t)$ satisfies the Schrödinger equation:

$$i\hbar \frac{\partial \psi(x,t)}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi(x,t)}{\partial x^2} + V(x) \psi(x,t)$$

Now we want to show $\psi^*(x,-t)$ is a solution as well.

PROOF:  (1) take complex conjugate

$$\Rightarrow -i\hbar \frac{\partial \psi^*(x,t)}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi^*(x,t)}{\partial x^2} + V(x) \psi^*(x,t)$$

[assuming the potential is time-independent $V(x,t) = V(x)$ here]

(2) change variable $t \rightarrow -t$

$$i\hbar \frac{\partial \psi^*(x,-t)}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi^*(x,-t)}{\partial x^2} + V(x) \psi^*(x,-t)$$

It's clear that $\psi^*(x,-t)$ also satisfies the Schrödinger equation.

Kramer's Degeneracy

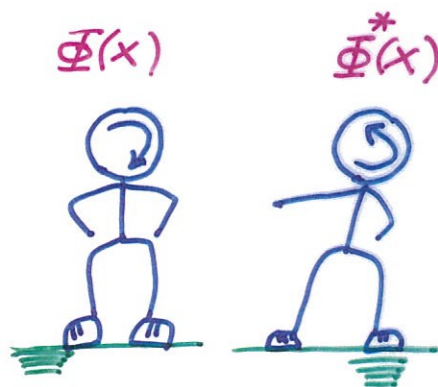
Suppose the wave function describes a stationary state with definite energy E :

$$\psi(x,t) = \Phi(x) e^{-i\omega t}$$

Its twin brother state under time reversal transformation is

$$\psi^*(x,-t) = \Phi^*(x) e^{i\omega(-t)} = \Phi^*(x) e^{-i\omega t} \quad \text{One notices that}$$

they share the same energy E !!



degenerate in energy !!

Example: plane-wave solution

$$\psi(x,t) = e^{ikx} e^{-i\omega t} \quad \text{R-moving}$$

$$\psi^*(x,-t) = e^{-ikx} e^{-i\omega t} \quad \text{L-moving}$$

★ By switching the direction of time, the $R \leftrightarrow L$ moving states also switch.

$E < 0$ Solution

Consider the $E < 0$ solution for a free particle.

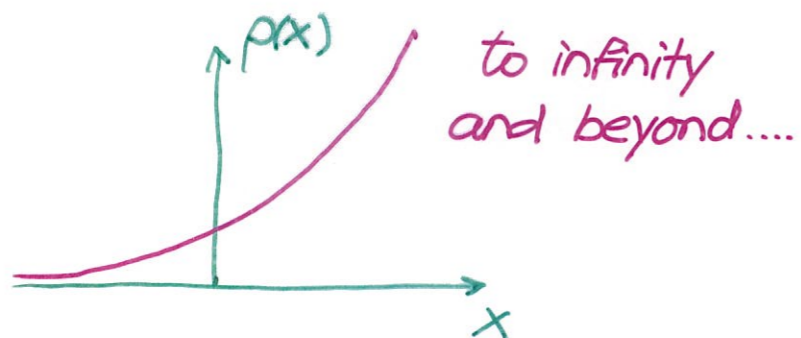
$$-\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} = -|E|\psi \quad \Rightarrow \quad \frac{d^2\psi}{dx^2} = \alpha^2 \psi \quad \alpha^2 = \frac{2m|E|}{\hbar^2}$$

The general solution can be written down rather easily,

$$\psi(x) = A e^{\alpha x} + B e^{-\alpha x} \quad \leftarrow \text{something wrong with this solution?? } \begin{matrix} \odot \\ \leftarrow \psi \\ \varepsilon \end{matrix}$$

Consider the special case where $B=0$

$$\psi(x) = A e^{\alpha x} \quad \rightarrow \quad \rho(x) = |\psi(x)|^2 = A^2 e^{2\alpha x}$$

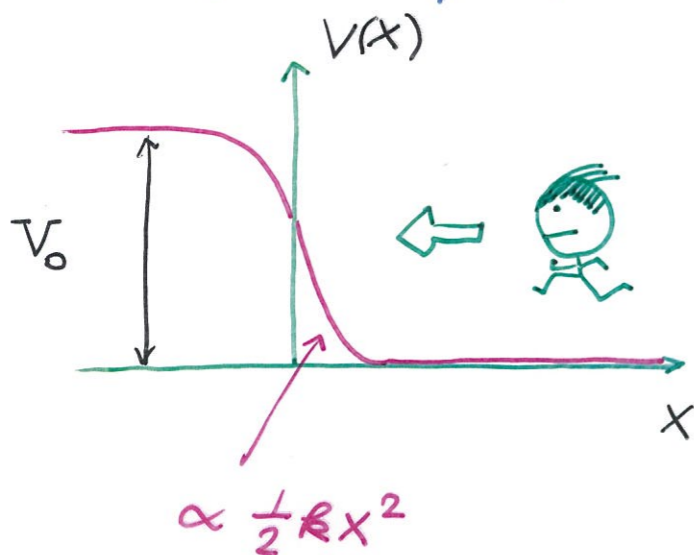


$$\int dx |\psi(x)|^2 \rightarrow \infty !!$$

The wave function diverges at $x \rightarrow +\infty$. Thus, the solution can not be used to generate a sensible probability density !!

Potential Wall.

Consider a particle and a wall at $x=0$. However, to simplify the problem, it is often approximated by the so-called hard-wall condition.



$$V(x) = \begin{cases} 0 & x > 0 \\ \infty & x \leq 0 \end{cases}$$

From linear superposition, the general solution with energy $E = \frac{(\hbar k)^2}{2m}$ takes the form:

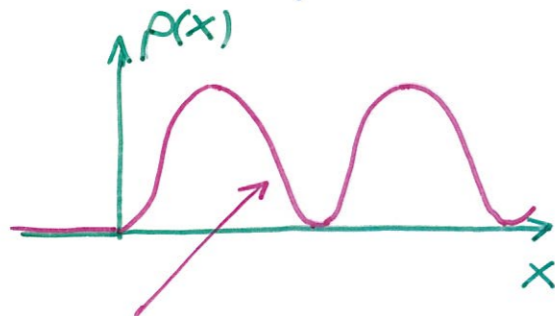
$$\psi(x) = A e^{ikx} + B e^{-ikx}$$

Since the wave function vanishes at $x=0$, we need to enforce the boundary condition $\psi(0) = 0$.

$$\psi(0) = 0 \Rightarrow A + B = 0 \quad B = -A.$$

$$\text{Thus, } \psi(x) = A e^{iRx} - A e^{-iRx} = 2iA \sin Rx = \underline{\underline{C \sin Rx}}$$

We can plot the probability density $P(x,t)$.



not uniform anymore!!

(1) $P(x,t)$ is not uniform anymore.

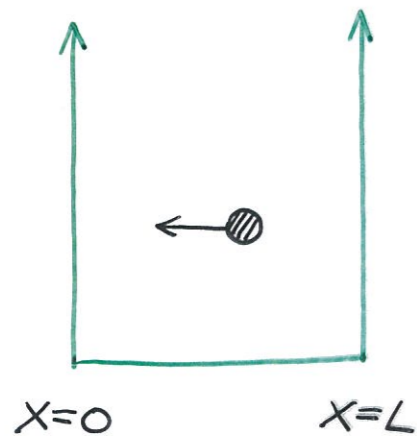
(2) $P(x,t)$ is static (not like the usual standing wave).

Furthermore, we can compute the probability current

$$j(x) = \frac{\hbar}{2mi} \left(\psi^* \frac{\partial \psi}{\partial x} - \frac{\partial \psi^*}{\partial x} \psi \right) = \frac{\hbar}{2mi} \left(\psi \frac{\partial \psi}{\partial x} - \frac{\partial \psi}{\partial x} \psi \right) = 0$$

- The probability current is zero as we expected.
- Hard wall gives rise to $k \rightarrow -k$ plus π shift !!

Potential Box



The Schrödinger eq reads $\frac{d^2\psi}{dx^2} + k^2\psi = 0$

where $k^2 = \frac{2mE}{\hbar^2}$. The general solution is

$$\psi(x) = A e^{ikx} + B e^{-ikx}$$

Now, needs to satisfy two boundary conditions

(1) $\psi(0) = 0$ (2) $\psi(L) = 0$ From the first B.C.

$$A + B = 0 \Rightarrow \psi(x) = A e^{ikx} - A e^{-ikx} = 2iA \sin kx \\ = C \sin kx$$

Now, ready to apply the 2nd B.C. $\psi(L) = 0$

$$\sin kL = 0 \Rightarrow k_n = \frac{n\pi}{L}$$

momentum is
quantized !!

Energy Quantization

From the quantized momentum $p_n = \hbar k_n = \frac{n\pi\hbar}{L}$, it's easy to show that energy is also quantized.

$$E_n = \frac{p_n^2}{2m} = \frac{n^2\pi^2\hbar^2}{2mL^2} \quad E_n \propto n^2 \text{ quantized!!}$$

We are NOT done yet.... $\psi_n(x) = C \sin(k_n x)$

Still need to figure out the const C $\ddot{\circ}$

$$\int dx |\psi(x)|^2 = 1 \quad \Rightarrow \quad \int_0^L dx C^2 \sin^2(k_n x) = 1$$

Recall that $\sin^2(k_n x) = \frac{1}{2} (1 + \cos(2k_n x))$

$$\int_0^L \sin^2(k_n x) dx = \int_0^L \frac{1}{2} + \frac{1}{2} \cos(2k_n x) dx = \frac{L}{2}$$

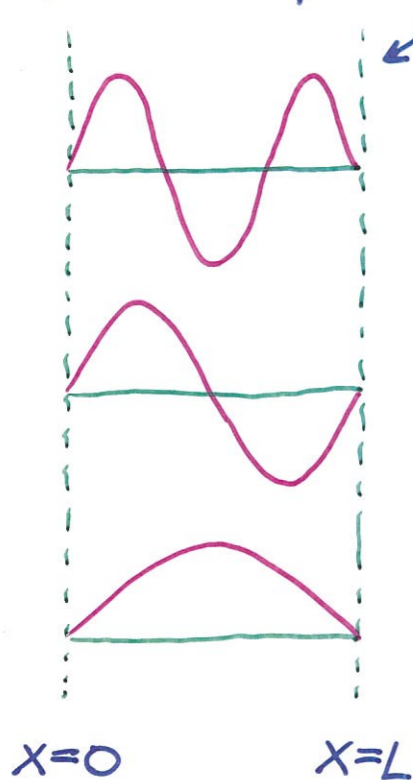
Thus, the wave function is $\psi_n(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right)$

Wave Function

Just as in conventional standing waves, we see nodal structure in wave functions. Note that, except the ground state, all excited states have nodes.

$$\# \text{ of nodes} = n - 1$$

probability amplitude

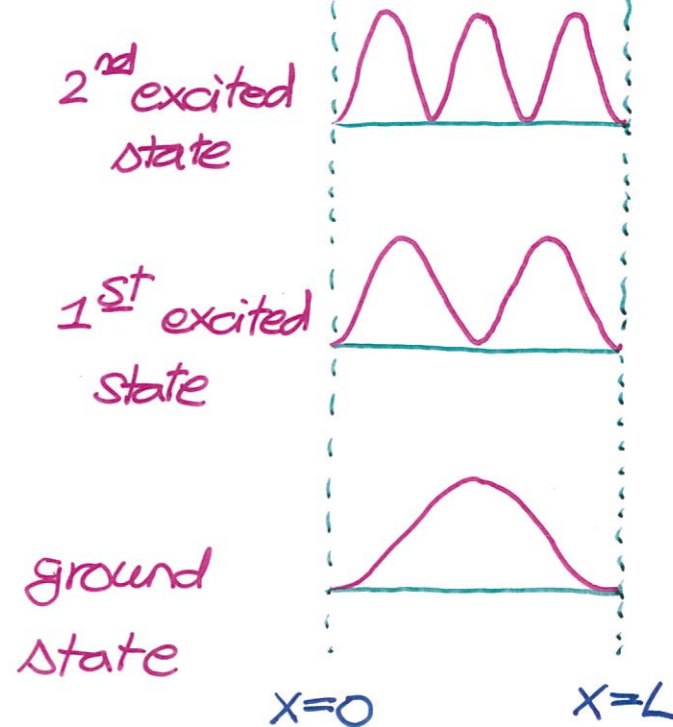


$$\psi_3(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{3\pi x}{L}\right)$$

$$\psi_2(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{2\pi x}{L}\right)$$

$$\psi_1(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{\pi x}{L}\right)$$

probability density

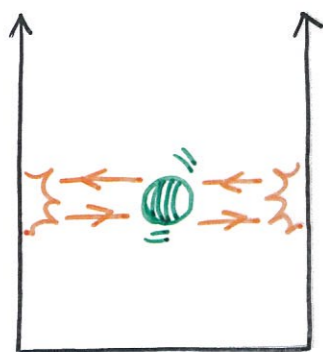


2nd excited state

1st excited state

ground state

Bohr Quantization



View the stationary solution as multiple self constructive interference. The phase difference

$$\text{is } \delta = 2\pi \left(\frac{2L}{\lambda} \right) - \varphi_s \quad \varphi_s = \pi + \pi$$

The constructive interference requires $\delta = 2n\pi$

$$2n\pi = 2\pi \left(\frac{2L}{\lambda} \right) - 2\pi \quad \Rightarrow \quad \frac{2L}{\lambda} = (n+1)$$

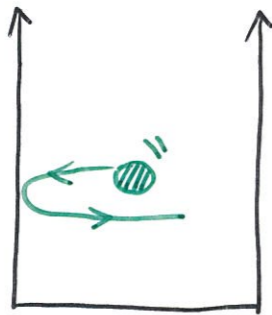
$$\text{Using the relation } p = \frac{h}{\lambda} \quad \Rightarrow \quad 2L \frac{h}{\lambda} = (n+1)h$$

Finally, we arrive at $2pL = n'h$. OR, in more formal format

$$\oint p dx = \hbar (2n\pi + \varphi_s)$$

↑
scattering phase.

Simple Applications.



$$\oint p dx = \hbar (2n\pi + \varphi_s) \Rightarrow p \cdot 2L = \hbar 2n\pi$$

$$p_n = \frac{2n\pi\hbar}{2L} = \frac{n\pi\hbar}{L}$$

Another example is Bohr's model for H atom. Somehow, the angular momentum is quantized....

$$\oint p dx = \hbar (2n\pi + \varphi_s)$$

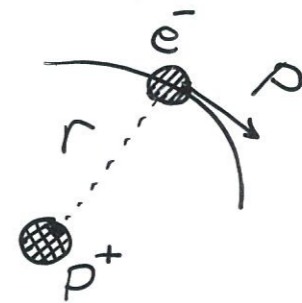
$\varphi_s = 0$ in this case !!

$$p \cdot 2\pi r = 2n\pi\hbar$$



$$L = n\hbar$$

the angular momentum is quantized in unit of \hbar .



Uncertainty Principle.

The wave function $\psi_n(x) = \sqrt{\frac{2}{L}} \sin(k_n x)$. We can calculate the uncertainty in position.

$$\langle x \rangle \equiv \int_0^L dx \rho(x) \cdot x = \frac{L}{2} \quad \text{by symmetry.}$$

$$\begin{aligned} (\Delta x)^2 &\equiv \langle (x - \langle x \rangle)^2 \rangle = \langle x^2 - 2x\langle x \rangle + \langle x \rangle^2 \rangle \\ &= \langle x^2 \rangle - 2\langle x \rangle\langle x \rangle + \langle x \rangle^2 = \langle x^2 \rangle - \langle x \rangle^2 \end{aligned}$$

Only need to compute $\langle x^2 \rangle$ now.

$$\langle x^2 \rangle = \int_0^L dx \rho(x) x^2 = \frac{2}{L} \int_0^L dx x^2 \sin^2(k_n x)$$

The integral is elementary:

$$\int x^2 \sin^2 kx dx = \frac{x^3}{6} - \frac{x \cos(2kx)}{4k^2} - \frac{2kx^2 - 1}{8k^3} \sin(2kx)$$

$$\langle x^2 \rangle = \int_0^L dx \frac{2}{L} \sin^2 k_n x \cdot x^2 = \frac{2}{L} \cdot \left[\frac{L^3}{6} - \frac{L}{4k_n^2} \right]$$

$$= \frac{1}{3} L^2 - \frac{1}{2k_n^2} \approx \frac{1}{3} L^2$$

Finally, the uncertainty in x is

$$(\Delta x)^2 = \langle x^2 \rangle - \langle x \rangle^2 \approx \frac{1}{3} L^2 - \frac{1}{4} L^2$$

$$= \frac{1}{12} L^2$$

⇒ $\Delta x \approx \frac{1}{2\sqrt{3}} L$

On the other hand, the uncertainty in momentum is

$$\langle p^2 \rangle = p_n^2$$

$$\langle p \rangle^2 = 0$$

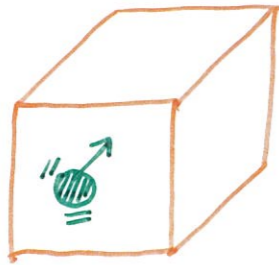
⇒ $\Delta p = \sqrt{\langle p^2 \rangle - \langle p \rangle^2} = p_n = \frac{n\hbar\pi}{L}$

One can easily see that $\Delta x \Delta p \approx \frac{1}{2\sqrt{3}} L \cdot \frac{n\hbar\pi}{L}$

→ $\Delta x \Delta p \approx n\hbar !! \quad \ddot{\sigma}$

Degeneracy and Symmetry

Now consider a particle in 3D box of length L . The wave function can be solved in a similar fashion.



$$\Psi(x, y, z) = \left(\frac{2}{L}\right)^{3/2} \sin \frac{n_x \pi x}{L} \sin \frac{n_y \pi y}{L} \sin \frac{n_z \pi z}{L}$$

The total energy $E = \frac{1}{2m} (P_x^2 + P_y^2 + P_z^2)$, plug in the quantized momenta P_x, P_y, P_z .

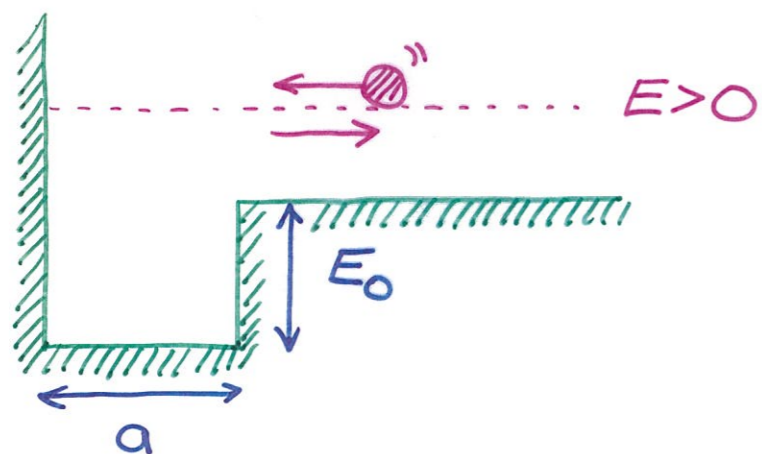
$$E = \frac{\hbar^2}{2m} (k_x^2 + k_y^2 + k_z^2) = \frac{\hbar^2 \pi^2}{2mL^2} (n_x^2 + n_y^2 + n_z^2)$$

The total energy depends on the quantum numbers (n_x, n_y, n_z)

$3E_1$	$(1, 1, 1)$	$D = 1$	related to symmetry.
$6E_1$	$(2, 1, 1), (1, 2, 1), (1, 1, 2)$	$D = 3$	
$9E_1$	$(2, 2, 1), (2, 1, 2), (1, 2, 2)$	$D = 3$	
\vdots	\vdots		

Finite Potential Well

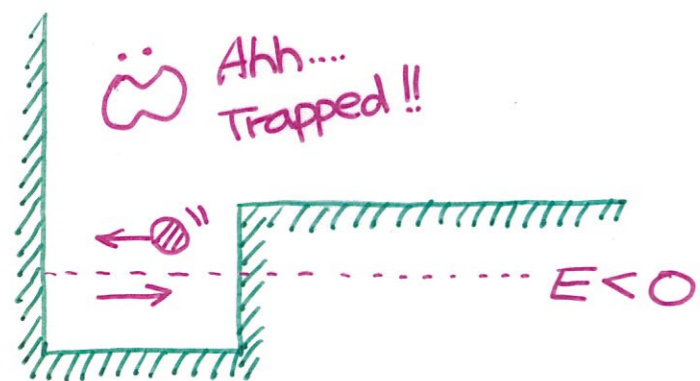
Consider a finite potential well as shown below. The



Schrödinger equation can be solved in two separated regimes easily. Then, we glue the solutions together.

In region I:

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi_I}{dx^2} - E_0 \psi_I = E \psi_I$$

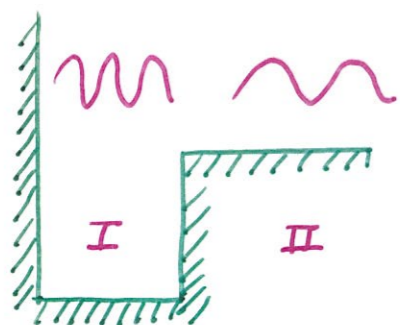


In region II:

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi_{II}}{dx^2} = E \psi_{II}$$

Both ψ_I and ψ_{II} can be solved easily. To glue them together smoothly, $\psi_I(a) = \psi_{II}(a)$, $\frac{d\psi_I}{dx}(a) = \frac{d\psi_{II}}{dx}(a)$

$E > 0$ Scattering state



Consider the positive energy solution. In both regimes, the solution is plane wave except the momentum is different.

In regime I:
$$-\frac{\hbar^2}{2m} \frac{d^2 \psi}{dx^2} - E_0 \psi = E \psi$$

The general solution is
$$\psi_I(x) = A e^{i k_I x} + B e^{-i k_I x}$$

where the momentum k_I is
$$\frac{\hbar^2 k_I^2}{2m} - E_0 = E \quad k_I = \sqrt{\frac{2m(E+E_0)}{\hbar^2}}$$

Due to the hard wall at $x=0 \Rightarrow \psi_I(x=0) = 0$

Thus, it requires $A+B=0$.

$$\psi_I(x) = A e^{i k_I x} - A e^{-i k_I x} = 2A i \sin k_I x = C \sin k_I x$$

We can now go ahead and compute $\psi_{II}(x)$ in regime II....

In regime II: $-\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} = E\psi \Rightarrow \psi_{II}(x) = A'e^{ik_{II}x} + B'e^{-ik_{II}x}$

It is equivalent to write the general solution in terms of sin, cos.

$$k_{II} = \sqrt{\frac{2mE}{\hbar^2}}$$

$$\psi_{II}(x) = D' \sin k_{II}x + E' \cos k_{II}x = c' \sin(k_{II}x + \delta)$$

Glue $\psi_I(x)$ and $\psi_{II}(x)$ together at $x=a$.

$$\psi_I(a) = \psi_{II}(a) \Rightarrow$$

$$c \sin(k_I a) = c' \sin(k_{II} a + \delta)$$

$$\frac{d\psi_I(a)}{dx} = \frac{d\psi_{II}(a)}{dx} \Rightarrow$$

$$c k_I \cos(k_I a) = c' k_{II} \cos(k_{II} a + \delta)$$

divide both eq's:

$$\frac{c \sin(k_I a)}{c k_I \cos(k_I a)} = \frac{c' \sin(k_{II} a + \delta)}{c' k_{II} \cos(k_{II} a + \delta)}$$

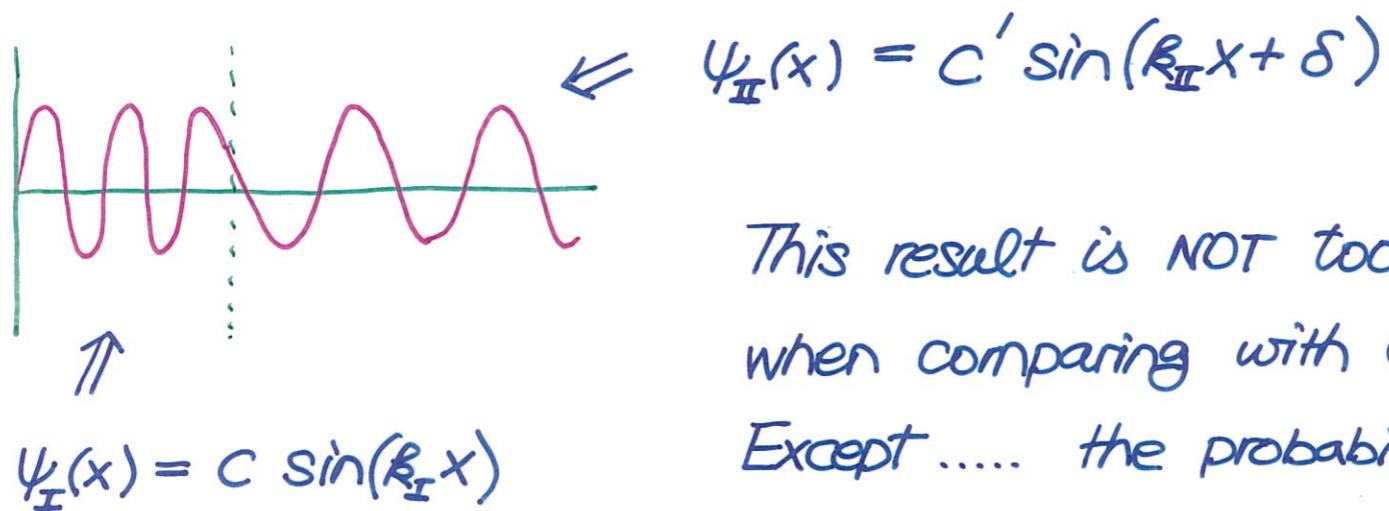
$$\Rightarrow \frac{1}{k_I} \tan(k_I a) = \frac{1}{k_{II}} \tan(k_{II} a + \delta)$$

★ δ can be solved !! ☺

Once the phase shift δ is known, the ratio between C and C' is

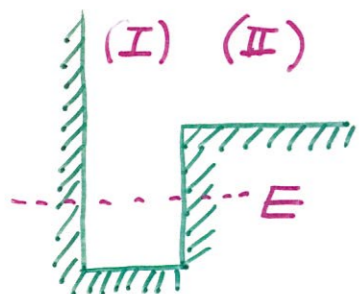
$$\frac{C'}{C} = \frac{\sin(k_I a)}{\sin(k_{II} a + \delta)}$$

Q: any missing equation??



This result is NOT too surprising when comparing with classical one. Except the probability density modulation in space \ddot{u}

$E < 0$ Bound State.



For $E < 0$ bound state, the solution inside the well is still plane wave. BUT! What about the solution in regime II? Let's take a closer look.

Following similar algebra, it is easy to show that

$$\psi_{\text{I}}(x) = C \sin(k_{\text{I}} x) \quad \text{where} \quad \frac{\hbar^2 k_{\text{I}}^2}{2m} - E_0 = -|E|$$

The momentum inside the well is $\hbar k_{\text{I}} = \sqrt{2m(E_0 - |E|)}$

Ok. Now turn our attention to regime II. The Schrödinger equation reads:

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi}{dx^2} = -|E| \psi \quad \Rightarrow \quad \frac{d^2 \psi}{dx^2} = \left(\frac{2m|E|}{\hbar^2} \right) \psi = \alpha^2 \psi$$

$$\psi_{\text{II}}(x) = \cancel{D e^{\alpha x}} + E e^{-\alpha x} \quad \text{with} \quad \hbar \alpha = \sqrt{2m|E|}$$

drop!! Why?

Collect the solution together:

$$\psi(x) = \begin{cases} C \sin k_I x & x \leq a \\ D e^{-\alpha x} & x \geq a \end{cases}$$

Try to glue them together
smooth and tight !!

We want (1) $\psi_I(a) = \psi_{II}(a)$ Need to do some math
(2) $\psi'_I(a) = \psi'_{II}(a)$ in the following $\vec{\mathcal{D}}$

From (1), $C \sin k_I a = D e^{-\alpha a} \Rightarrow k_I \cot k_I a = -\alpha$

From (2), $C k_I \cos k_I a = -D \alpha e^{-\alpha a}$

Writing the absolute value of energy $|E| = \mathcal{E}$

$$\hbar k_I = \sqrt{2m(E_0 - \mathcal{E})} \Rightarrow \hbar^2 k_I^2 = 2mE_0 - 2m\mathcal{E}$$

$$\hbar \alpha = \sqrt{2m\mathcal{E}} \Rightarrow \hbar^2 \alpha^2 = 2m\mathcal{E}$$

$$\alpha = \sqrt{\frac{2mE_0}{\hbar^2} - k_I^2}$$

The matching B.C. gives $k_I \cot(k_I a) = -\alpha$

$$-\cot(k_I a) = \frac{\alpha}{k_I} \quad \Rightarrow \quad -\cot(k_I a) = \frac{\sqrt{\frac{2mE_0}{\hbar^2} - k_I^2}}{k_I}$$

The bound state energy can be solved by plotting both sides and looking for intersections. Easier if we make everything dimensionless.

$$y = k_I a$$
$$\lambda = \frac{2mV_0 a^2}{\hbar^2}$$

$$-\cot y = \frac{\sqrt{\lambda - y^2}}{y}$$

↙ energy is quantized !!

One can read off the minimum λ for the bound state to exist.

$$\lambda - \left(\frac{\pi}{2}\right)^2 \geq 0 \quad \lambda \geq \frac{\pi^2}{4}$$

Quantum Leakage.

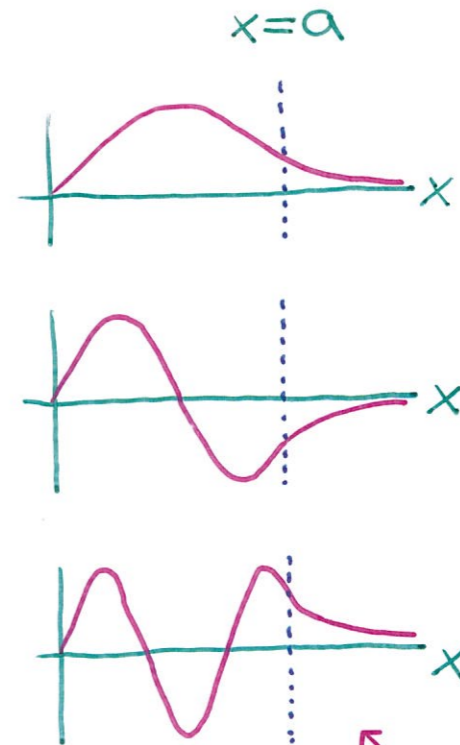
From the matching B.C.

$$-\cot(k_I a) = \frac{\alpha}{k_I}$$

$$\alpha = \sqrt{\frac{2mE_0}{\hbar^2} - k_I^2}$$

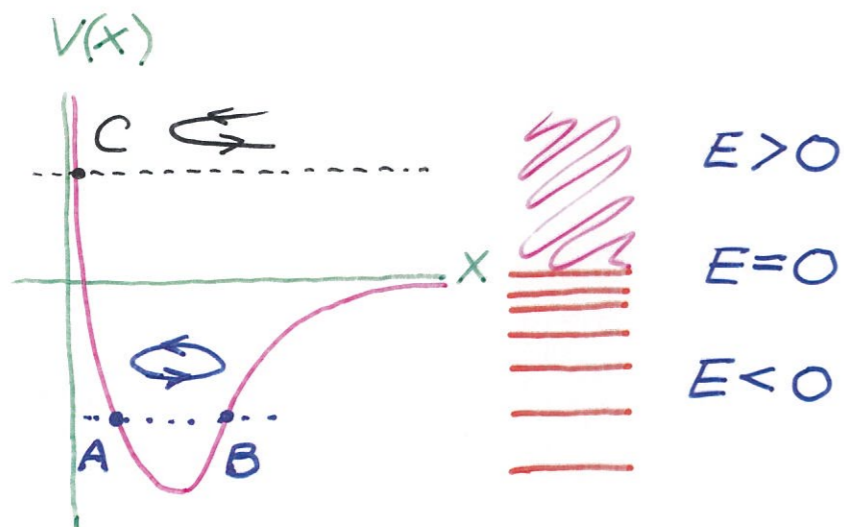
For steeper potential well, the number of bound states increase. Several important key features:

- (1) Nodal Structure
- (2) Quantum Leakage: it is possible to find the particle in the classically forbidden regime !!
- (3) The decaying solution does exist !!



$x=a$
forbidden
regime by
classical mechanics.

Particle in General Potential



Consider a particle moving in the potential (shown in the left).

Classically 

(1) Bounded motion for $E < 0$.
The particle moving between two turning points.

(2) Unbounded motion for $E \geq 0$

The particle moves inward until the turning point. Then, it changes direction and moves out to infinity.

Turning Point :



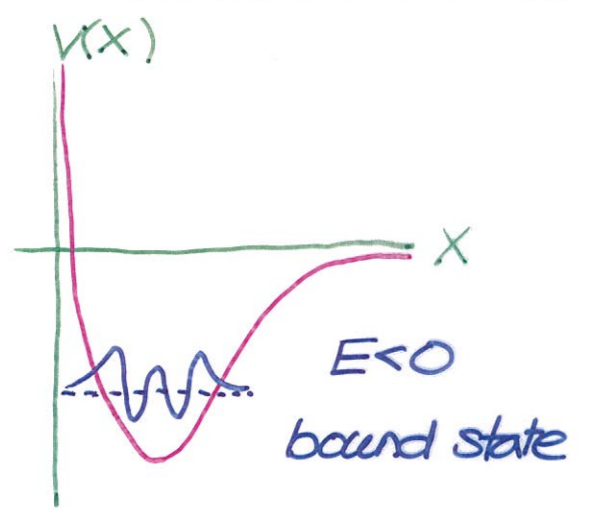
$$\frac{p^2}{2m} + V(x) = E$$

momentum is zero

$$\Rightarrow V(x_c) = E$$

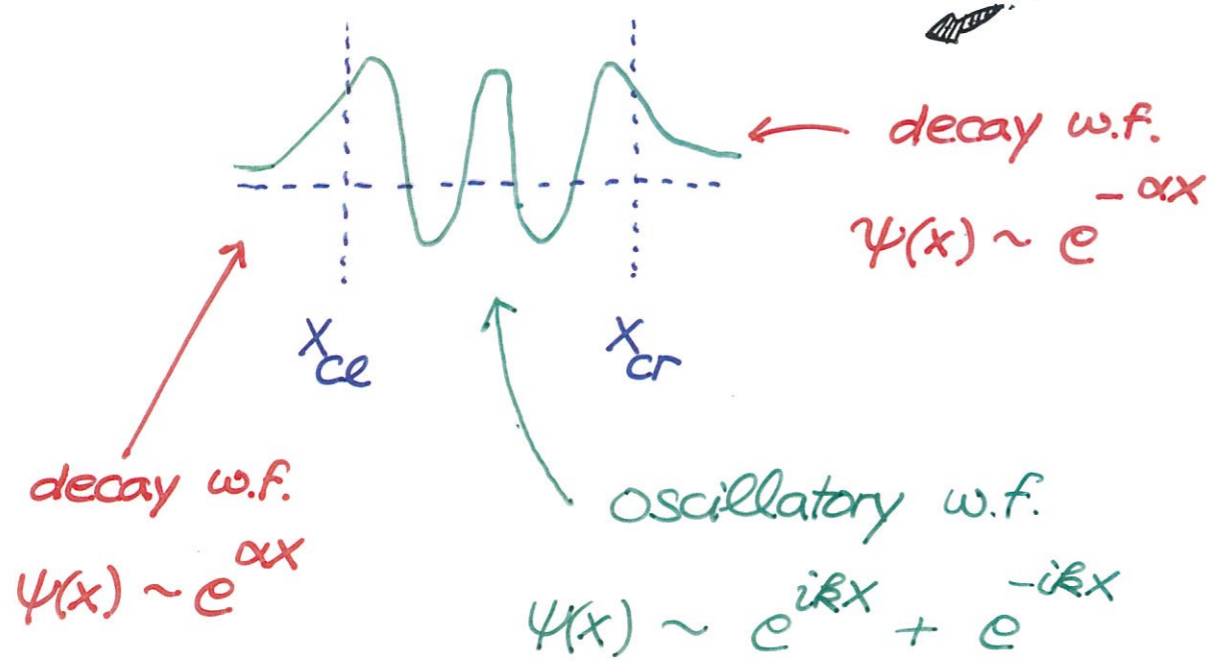
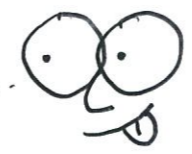
- 1 turning point for open orbits.
- 2 turning points for closed

What's new is quantum physics ??



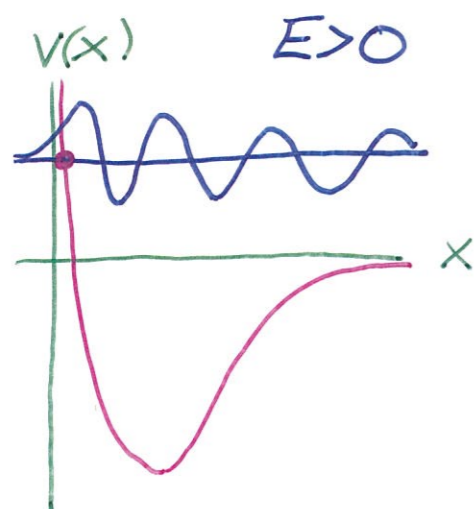
(1) For $E < 0$, bound state with discrete energy levels. The typical wave fn is shown on the left.

Look closer into the details of the wave function



Solving for the exact wave function can be a true challenge... BUT !! It's quite easy to get a rough idea about the shape of wave function.

Continuum



(2) For $E > 0$, continuum with **continuous** energy levels. Again, one can get a rough idea about wave function easily.

Observation: 

$$\frac{p^2}{2m} + V(x) = E \Rightarrow \text{position-dependent momentum}$$

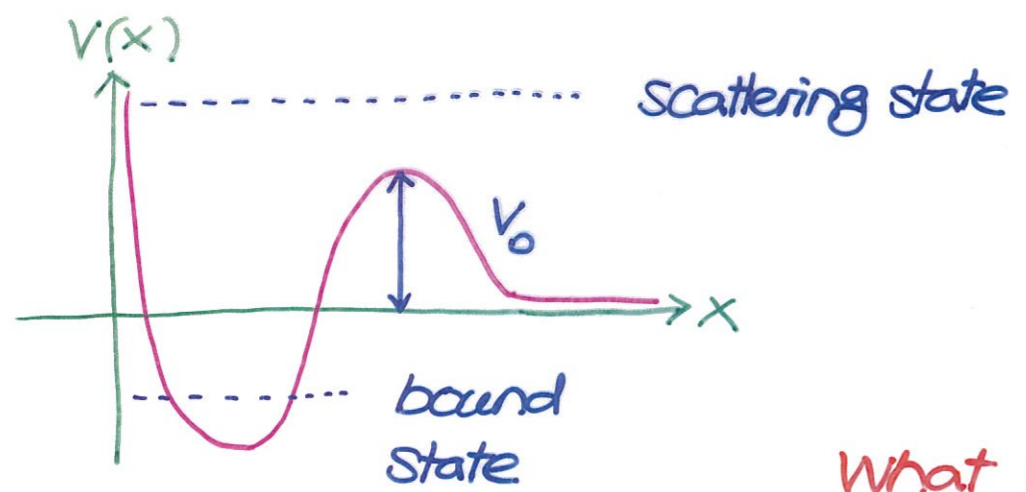
$$p(x) = \pm \sqrt{2m[E - V(x)]}$$

Thus, we can construct the wave function:

$$\psi(x) \sim A e^{iR(x)x} + B e^{-iR(x)x} = A' \sin[R(x)x + \delta]$$

where the position-dep $p(x) = \hbar R(x)$ and δ is the phase shift.

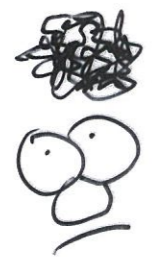
Quantum Tunneling



Consider the potential profile.
It's clear that ...

- (1) $E < 0$ bound state
- (2) $E > V_0$ scattering state

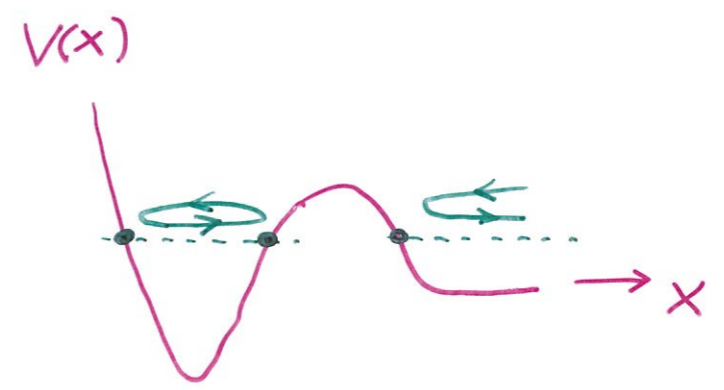
What about the range $0 < E < V_0$?



Think, think, think ... Classically, we still find either

- (1) 2 turning points, OR, (2) 1 t.p.

Depending on the initial condition, the particle can be inside the trap with closed orbit, or outside the trap with open orbit !!



Again, what about quantum mechanics?

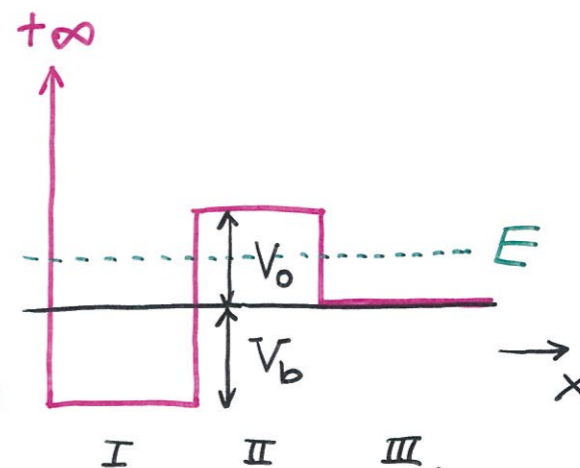


Let's simplify the question a bit to
We can write down the
wave function easily.

$$\psi_I(x) = A e^{i\beta x} + B e^{-i\beta x}$$

$$\psi_{II}(x) = C e^{-\alpha x} + D e^{\alpha x} \quad \leftarrow \text{quantum surprise!!}$$

$$\psi_{III}(x) = E e^{i\beta' x} + F e^{-i\beta' x}$$



The momentum in regimes I & III is

$$\frac{p^2}{2m} + (-V_b) = E \quad p = \sqrt{2m(E+V_b)} \quad \text{regime I.}$$

$$\frac{p'^2}{2m} + 0 = E \quad p' = \sqrt{2mE} \quad \text{regime III.}$$

In regime II, the "negative energy" state is

$$\frac{p^2}{2m} + V_0 = E \Rightarrow \frac{p^2}{2m} = E - V_0 < 0 \quad !!$$

$$\frac{p^2}{2m} = -(V_0 - E), \quad p = \pm i \sqrt{2m(V_0 - E)}$$

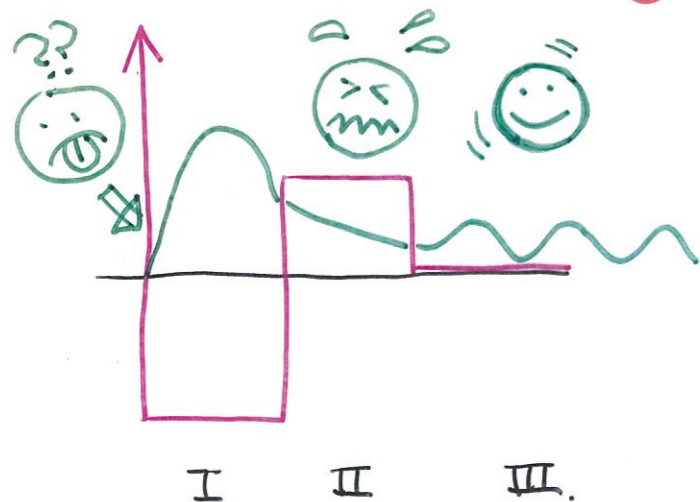
Thus, the decaying parameter $\alpha = ip = \pm \sqrt{2m(V_0 - E)}$

$$\psi_{\text{II}}(x) = C e^{ipx} + D e^{-ipx} = C e^{-\alpha x} + D e^{\alpha x}$$

★ Note that, it is impossible to find the particle in regime II in *classical limit*. However, in quantum mechanics, an *imaginary momentum* is meaningful and gives rise to the *spatially decaying solutions* !!

★ The presence of these decaying states give quantum tunneling !!

Quantum Tunneling !!



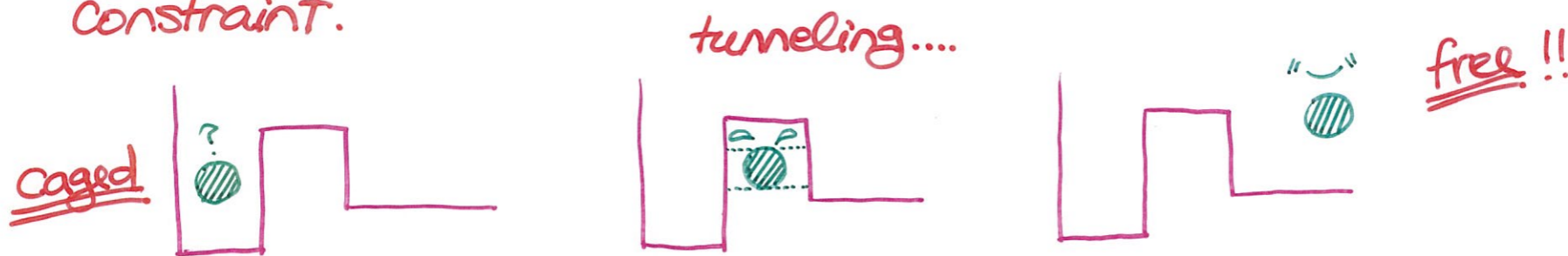
Let's look at the wave function again.

$$\Psi_I(x) = A e^{i\kappa x} + B e^{-i\kappa x}$$

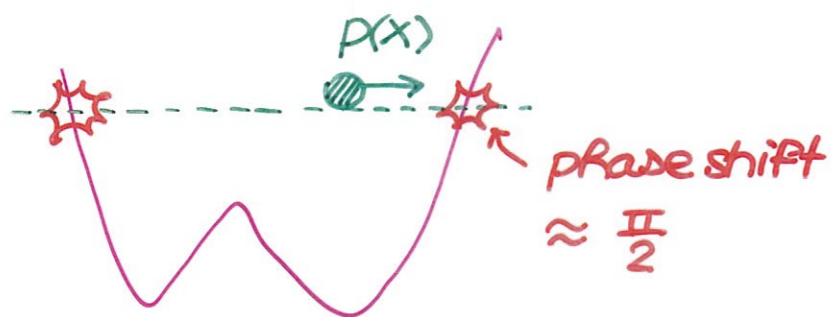
$$\Psi_{II}(x) = C e^{-\alpha x} + D e^{\alpha x}$$

$$\Psi_{III}(x) = E e^{i\kappa' x} + F e^{-i\kappa' x}$$

By matching the boundary conditions, $(\Psi, \frac{d\Psi}{dx})$ must be continuous \Rightarrow 4 constraints. plus 1 constraint at $x=0$ [$\Psi_I(0) = 0$]. We can solve the constants. Remember the normalization condition gives the final 1 constraint.



Semiclassical Approximation

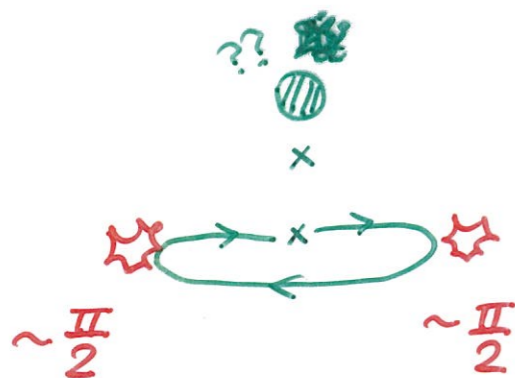


The first cute notion is $p(x)$ — the position-dependent momentum.

$$\frac{p^2}{2m} + V(x) = E, \quad \text{smooth } V(x).$$

$$p^2(x) = 2m[E - V(x)] > 0 \quad \Rightarrow \quad p(x) = \pm \sqrt{2m[E - V(x)]}$$

The second cute notion is **constructive self interference**.



$$\frac{1}{\hbar} \oint p(x) dx - \varphi_s = 2n\pi, \quad \varphi_s = \text{scattering phase}$$

$$\frac{1}{\hbar} \oint p(x) dx = 2n\pi + \varphi_s$$

$$\oint \sqrt{2m[E - V(x)]} dx = \hbar (2n\pi + \varphi_s)$$

$$\Rightarrow \oint \sqrt{2m[E - V(x)]} = 2\pi\hbar \left(n + \frac{1}{2}\right)$$

A simple check.....

Does the semiclassical approximation make sense at all?

Let's check.... The position-dependent momentum $p(x)$ implies the wave function takes the form,

$$\psi(x,t) \approx e^{i[\mathcal{R}(x)x - \omega t]} \quad p(x) = \hbar \mathcal{R}(x) = \sqrt{2m(E-V)}$$

check. ✓ $\frac{\partial \psi}{\partial t} = (-i\omega) e^{i(\mathcal{R}x - \omega t)} = (-i\omega) \psi$

$$\frac{\partial \psi}{\partial x} = i \left(\mathcal{R} + \frac{d\mathcal{R}}{dx} x \right) e^{i(\mathcal{R}x - \omega t)} = i \left(\mathcal{R} + \frac{d\mathcal{R}}{dx} x \right) \psi$$

$$\begin{aligned} \frac{\partial^2 \psi}{\partial x^2} &= i \left(\frac{d\mathcal{R}}{dx} + \frac{d^2\mathcal{R}}{dx^2} x + \frac{d\mathcal{R}}{dx} \right) e^{i(\mathcal{R}x - \omega t)} + (i)^2 \left(\mathcal{R} + \frac{d\mathcal{R}}{dx} x \right)^2 e^{i(\mathcal{R}x - \omega t)} \\ &= \left[- \left(\mathcal{R} + \frac{d\mathcal{R}}{dx} x \right)^2 + i \left(2 \frac{d\mathcal{R}}{dx} + \frac{d^2\mathcal{R}}{dx^2} x \right) \right] \psi \end{aligned}$$

Note that $\frac{d\mathcal{R}}{dx} = \frac{1}{\hbar} \frac{dp}{dx} = \frac{1}{\hbar} \frac{-2m \frac{dV}{dx}}{2\sqrt{2m(E-V)}} \propto \frac{dV}{dx} !!$ Thus, for

smooth potential, we can drop all derivatives $\ddot{}$

After dropping all spatial derivatives

$$\frac{\partial \psi}{\partial t} = (-i\omega) \psi$$

$$\frac{\partial^2 \psi}{\partial x^2} \approx -k^2 \psi$$

plug in the
Schrödinger
equation



$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V\psi$$

$$\text{left} = i\hbar \frac{\partial \psi}{\partial t} = (i\hbar)(-i\omega) \psi = \hbar\omega \psi = \underline{E \psi}$$

$$\text{right} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V\psi = \left[-\frac{\hbar^2}{2m} (-k^2) + V \right] \psi = \left[\frac{p^2}{2m} + V \right] \psi$$

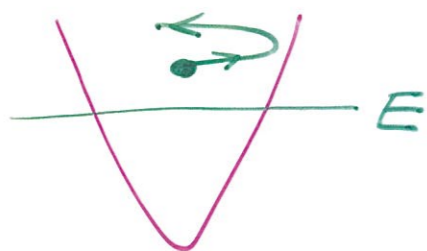
$$= \left[\frac{(\sqrt{2m(E-V)})^2}{2m} + V \right] \psi = [(E-V) + V] = \underline{E \psi}$$

The left = right \Rightarrow Schrödinger equation is satisfied.

Thus, as long as the potential is **smooth**, one can use the semiclassical approximation to compute the discrete bound-state energy.

Q: How "smooth" is smooth? Huh? 

Try a simple example.



Consider a parabolic potential $V(x) = \frac{1}{2}kx^2$.

The Schrödinger equation is

$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + \frac{1}{2}kx^2 \psi = E \psi$$

The above differential equation certainly looks unfriendly.
At least, to freshman....

Let's try to solve it by the semiclassical approach.

$$\oint p(x) dx = \hbar (2n\pi + \phi_s) \Leftrightarrow \oint \sqrt{2m(E-V)} dx = 2\pi\hbar (n + \frac{1}{2})$$

How nice!! We only need to compute a simple integral.

$$\oint p(x) dx = 2 \int_{-x_c}^{x_c} \sqrt{2mE - mkx^2} dx, \quad \pm x_c: \text{turning points}$$

Some algebra....

$$2 \int_{-x_c}^{x_c} \sqrt{2mE - mRx^2} dx = 2\sqrt{mR} \int_{-x_c}^{x_c} \sqrt{\frac{2E}{R} - x^2} dx$$

$$\frac{1}{2}Rx_c^2 = E$$

$$x_c^2 = \frac{2E}{R}$$

$$= 2\sqrt{mR} \int_{-x_c}^{x_c} \sqrt{x_c^2 - x^2} dx$$

change variable θ

$$x = x_c \sin \theta$$

$$\begin{cases} x = x_c \rightarrow \theta = \frac{\pi}{2} \\ x = -x_c \rightarrow \theta = -\frac{\pi}{2} \end{cases}$$

$$dx = x_c \cos \theta d\theta$$

$$= 2\sqrt{mR} \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} x_c \sqrt{1 - \sin^2 \theta} (x_c \cos \theta d\theta)$$

$$= 2\sqrt{mR} x_c^2 \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \cos^2 \theta d\theta = 2\sqrt{mR} \frac{2E}{R} \cdot (\text{integral})$$

$$= \sqrt{\frac{m}{R}} \cdot 4E \cdot (\text{integral}) = \frac{4E}{\omega} \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \cos^2 \theta d\theta$$

Now the integral... 

a piece of cake

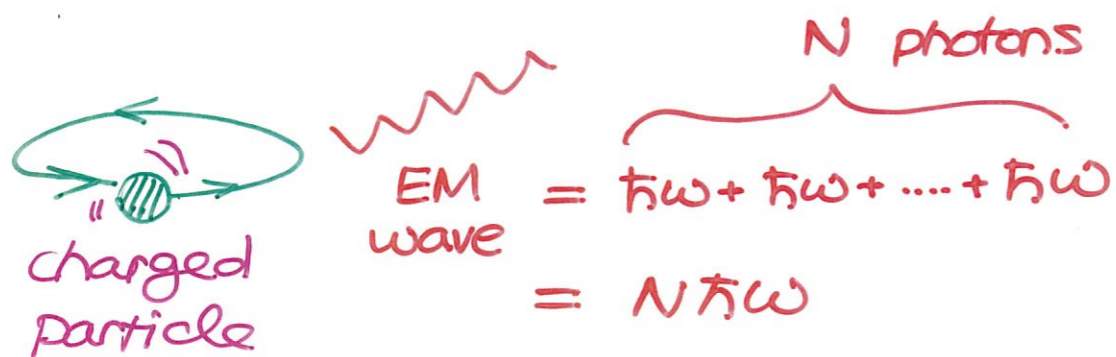


$$\int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \cos^2 \theta d\theta = \int_{-\frac{\pi}{2}}^{\frac{\pi}{2}} \frac{1 + \cos 2\theta}{2} d\theta = \frac{\theta}{2} \Big|_{-\frac{\pi}{2}}^{\frac{\pi}{2}} + \frac{1}{4} \sin 2\theta \Big|_{-\frac{\pi}{2}}^{\frac{\pi}{2}} = \frac{\pi}{2}$$

Finally, combine every pieces together....

$$\frac{4E}{\omega} \cdot \frac{\pi}{2} = 2\pi\hbar \left(n + \frac{1}{2}\right) \Rightarrow E_n = \hbar\omega \left(n + \frac{1}{2}\right)$$

★ For a simple harmonic oscillator, the "allowed" energy is no longer continuous. In fact, the energy is quantized in unit of $\hbar\omega = h\nu$ - very similar to Einstein's notion for photons !!



★ The residual $\frac{1}{2}$ comes from
 - Heisenberg uncertainty principle !!



THE END