

HH0060 Partition Function

Starting from the Boltzmann factor, $P(\epsilon_s) \propto e^{-\epsilon_s/\tau}$, it is convenient to introduce the partition function Z ,

$$Z = \sum_s e^{-\epsilon_s/\tau} \quad \rightarrow \quad P(\epsilon_s) = \frac{1}{Z} e^{-\epsilon_s/\tau}$$

very very useful ☺

The average energy of the system can be computed from the partition function.

$$U = \langle \epsilon \rangle = \sum_s \epsilon_s P(\epsilon_s) = \frac{1}{Z} \sum_s \epsilon_s e^{-\epsilon_s/\tau}$$

Compare both expressions.

Notice that $\frac{\partial Z}{\partial \tau} = \sum_s e^{-\epsilon_s/\tau} \cdot (-\epsilon_s) \left(-\frac{1}{\tau^2}\right) = \frac{1}{\tau^2} \sum_s \epsilon_s e^{-\epsilon_s/\tau}$

One obtains the very useful thermal identity,

$$U = -\tau^2 \frac{\partial \log Z}{\partial \tau}$$

Let's apply these results to a binary spin.

① A binary spin in thermal equilibrium

There are two states $\epsilon_s = \pm mB = \pm \epsilon$ for a binary spin. The partition function is rather simple,



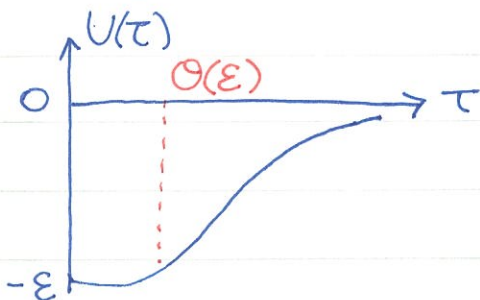
$$Z = e^{\epsilon/\tau} + e^{-\epsilon/\tau} = 2 \cosh(\epsilon/\tau)$$

Taking derivative with respect to $\tau \rightarrow$ average energy U

$$U = -\tau^2 \frac{1}{Z} \frac{\partial Z}{\partial \tau} = -\tau^2 \cdot \frac{1}{2 \cosh(\epsilon/\tau)} \cdot 2 \sinh(\epsilon/\tau) \cdot \left(-\frac{\epsilon}{\tau^2}\right)$$

$$\rightarrow \quad U = -\epsilon \tanh(\epsilon/\tau)$$

$\tau \rightarrow 0, U \rightarrow -\epsilon$ ✓ Note that U
 $\tau \rightarrow \infty, U \rightarrow 0$ ✓ is always negative



It shall be clear that $U \approx -\epsilon$ for $\tau \ll \epsilon$ where the ground state reigns. On the other hand, for $\tau \gg \epsilon$, thermal fluctuations dominates and $U \approx \frac{1}{2}(-\epsilon + \epsilon) = 0$!

According to the first law $dU = \tau d\sigma - p dV + \mu dN$

At constant $V, N,$

$dU = \tau d\sigma \rightarrow$ Introduce the heat capacity C_V at const. $V.$

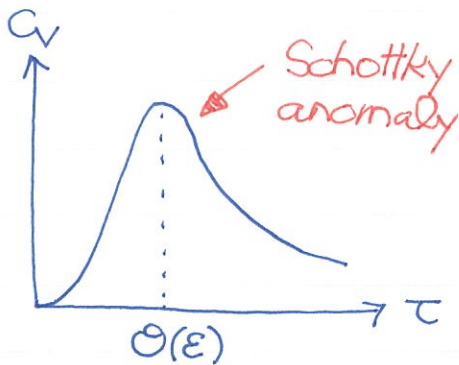
$$C_V = \tau \left(\frac{\partial \sigma}{\partial \tau} \right)_V = \left(\frac{\partial U}{\partial \tau} \right)_V, \text{ note that } \tau d\sigma = \text{infinitesimal heat.}$$

$$C_V = \left(\frac{\partial U}{\partial \tau} \right)_V = -\varepsilon \operatorname{sech}^2(\varepsilon/\tau) \cdot \left(-\frac{\varepsilon}{\tau^2} \right) = \left(\frac{\varepsilon}{\tau} \right)^2 \frac{1}{\cosh^2(\varepsilon/\tau)}$$

For $\tau \ll \varepsilon,$ $\cosh(\varepsilon/\tau) \approx \frac{1}{2} e^{\varepsilon/\tau},$ $C_V \approx \left(\frac{2\varepsilon}{\tau} \right)^2 e^{-2\varepsilon/\tau}$

The heat capacity is exponentially small at low temperatures. The exponential suppression arises from the energy gap $\Delta = 2\varepsilon$ in the binary spin system.

For $\tau \gg \varepsilon,$ $\cosh(\varepsilon/\tau) \approx 1,$ $C_V \approx \left(\frac{\varepsilon}{\tau} \right)^2$



The heat capacity is small for both high τ and low τ limits. But, there is a hump when $\tau \sim \varepsilon.$ Why?

From the definition,

$$\Delta\sigma = \int_0^\tau \frac{C_V}{\tau'} d\tau', \quad \Delta\sigma = \sigma(\tau) - \sigma(0)$$

For a binary spin, we know $\sigma(0) = \log(g(\tau=0)) = 0.$

$$\sigma(\tau) = \int_0^\tau \frac{C_V}{\tau'} d\tau'$$

Introduce the dimensionless variable

$$x = \tau/\varepsilon \rightarrow C_V = \frac{1}{x^2} \frac{1}{\cosh^2(1/x)}$$

$$\sigma(\tau) = \int_0^{\tau/\varepsilon} \frac{C_V}{x} dx = \int_0^{\tau/\varepsilon} \frac{1}{x^3} \frac{1}{\cosh^2(1/x)} dx \quad \leftarrow \text{look up the integral table.}$$

$$= \log\left[\cosh\left(\frac{1}{x}\right)\right] - \frac{1}{x} \tanh\left(\frac{1}{x}\right) - [-\log 2]$$

OK, OK, I lied.... I used Mathematica for integrals ☹

It is rather nice that the entropy $\sigma(\tau)$ obtained this way is identical to Shannon entropy.

$$\sigma_I \equiv \sum_s -P_s \log P_s = \langle -\log P \rangle \quad P_s = \frac{1}{Z} e^{-\epsilon_s/\tau}$$

$$= \langle \log Z + \epsilon/\tau \rangle \Rightarrow$$

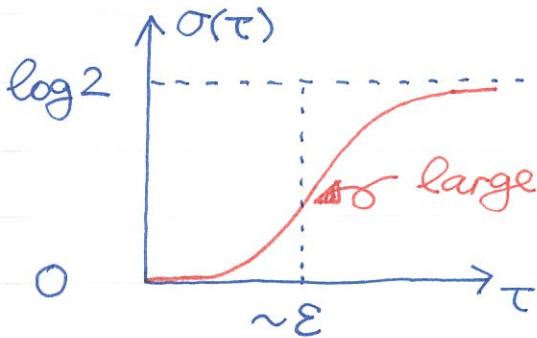
$$\sigma_I = \log Z + U/\tau$$

entropy for Boltzmann distribution.

We already calculate Z and U before,

$$\sigma_I = \log [2 \cosh(\epsilon/\tau)] - (\frac{\epsilon}{\tau}) \tanh(\frac{\epsilon}{\tau})$$

$$= \log [\cosh(\frac{1}{x})] - \frac{1}{x} \tanh(\frac{1}{x}) + \log 2 = \sigma(\tau)$$



① For $\tau \ll \epsilon$, $P(-\epsilon) = 1$, $P(\epsilon) \approx 0$, the entropy is almost 0.

② For $\tau \gg \epsilon$, $P(-\epsilon) \approx P(\epsilon) \approx \frac{1}{2}$, the entropy is almost $\log 2$.

Thus, in both low temp. and high

temp. limits, the entropy does not change much. It means that heat capacity C_V is small in both limits.

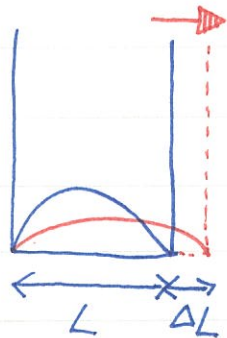
Only around $\tau \sim \epsilon$, entropy increases significantly, leading to the Schottky anomaly in C_V .

② Pressure: Consider a particle in a 1D box. The energy is

$$\epsilon_n = \frac{\hbar^2 k_n^2}{2m} = \frac{\hbar^2 \pi^2 n^2}{2mL^2} \propto \frac{1}{L^2}$$

Changing L gives rise to energy change,

$$\Delta \epsilon_n = \frac{\partial \epsilon_n}{\partial L} \cdot \Delta L \xrightarrow{3D} \frac{\partial \epsilon_n}{\partial V} \Delta V, \quad \Delta V = A \Delta L$$



Assume the probability distribution does not change during volume expansion (constant σ). The average

energy of the system also changes,

$$\Delta U = \langle \Delta \mathcal{E} \rangle = \left\langle \frac{\partial \mathcal{E}}{\partial V} \right\rangle \Delta V. \quad \text{For positive pressure, } \frac{\partial \mathcal{E}_n}{\partial V} < 0$$

and vice versa. Thus, we define pressure p as

$$p \equiv - \left\langle \frac{\partial \mathcal{E}}{\partial V} \right\rangle = - \sum_n P_n \frac{\partial \mathcal{E}_n}{\partial V} = - \frac{\partial}{\partial V} \sum_n P_n \mathcal{E}_n = - \left(\frac{\partial U}{\partial V} \right)_\sigma$$

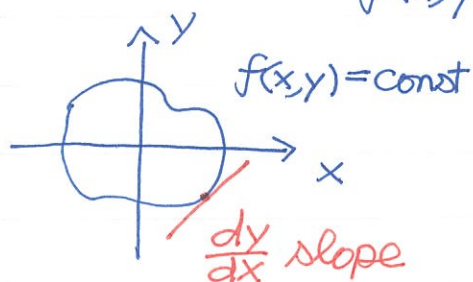
Finally, we obtain the important relation

$$p = - \left(\frac{\partial U}{\partial V} \right)_\sigma$$

Now, we would like to show another expression for p .

review.

$$f(x, y) = \text{const} \rightarrow df = 0, \quad \frac{\partial f}{\partial x} dx + \frac{\partial f}{\partial y} dy = 0$$



$$\text{Thus, } \frac{dy}{dx} = - \frac{\frac{\partial f}{\partial x}}{\frac{\partial f}{\partial y}} \quad \text{simple to}$$

Here comes the key!

$$\frac{dy}{dx} = \left(\frac{\partial y}{\partial x} \right)_f \quad \text{This is the proper notation!}$$

Now make the following replacement, $f \rightarrow \sigma$, $x \rightarrow V$, $y \rightarrow U$

$$\left(\frac{\partial U}{\partial V} \right)_\sigma = - \frac{\left(\frac{\partial \sigma}{\partial V} \right)_U}{\left(\frac{\partial \sigma}{\partial U} \right)_V} \quad \text{note that } \left(\frac{\partial \sigma}{\partial U} \right)_V = \frac{1}{\tau} \quad \text{and } \left(\frac{\partial U}{\partial V} \right)_\sigma = -p$$

We arrive at a non-trivial expression

$$p = \tau \left(\frac{\partial \sigma}{\partial V} \right)_U$$



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2012.10.10